

Družicové metody v geodézii a katastru
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**GRAIL and LOLA satellite data resolve
the long-lasting convergence/divergence problem
for the analytical downward continuation
of the external spherical harmonic expansions**

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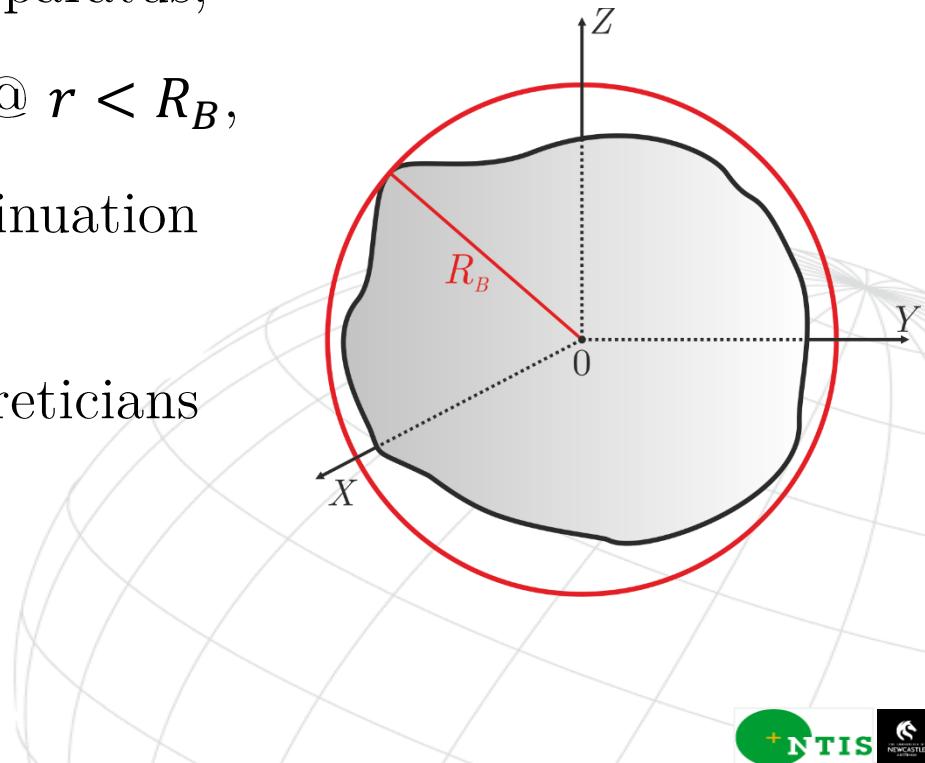
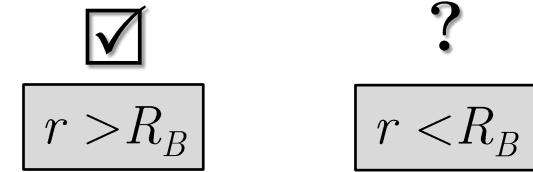
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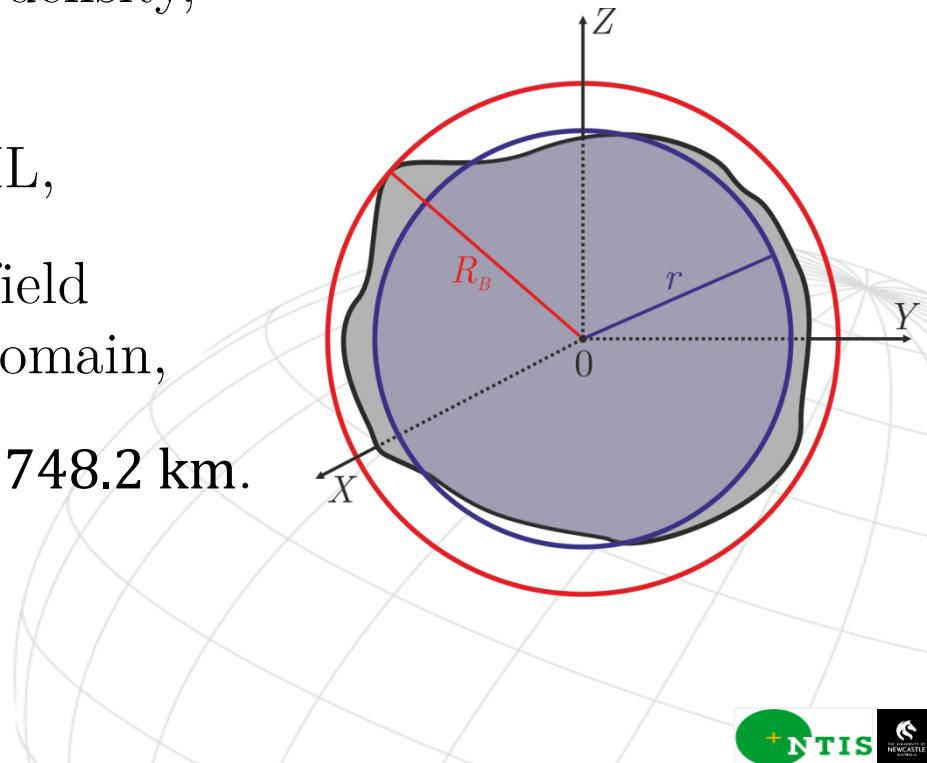
1. Introduction:

- External Spherical Harmonic Expansions (SHEs) – popular mathematical apparatus,
- Converge $\forall r > R_B$, but employed @ $r < R_B$,
- Justification of the analytical continuation questionable,
- Contradictory conclusions by theoreticians and pragmatists.



2. This presentation:

- Lunar gravitational field inferred by topographic masses of constant density,
- Assumption: perfect correlation with observations, e.g., from GRAIL,
- Investigation of two gravitational field quantities in spectral and spatial domain,
- Analysis @ $r = 1738 \text{ km} < R_B = 1748.2 \text{ km}$.



3. First-order radial derivative of V :

External SHE:

$$V_r(r, \Omega) = -\frac{GM}{R^2} \sum_{n,m}^N \left(\frac{R}{r}\right)^{n+2} (n+1) \bar{C}_{n,m} \bar{Y}_{n,m}(\Omega)$$

Internal SHE:

$$V_r(r, \Omega) = -\frac{GM}{R_o^2} \sum_{n,m}^N \left(\frac{R_o}{r}\right)^{n+2} (n+1) \bar{C}_{n,m}^o(r) \bar{Y}_{n,m}(\Omega)$$

$$+ \frac{GM}{R_i^2} \sum_{n,m}^N \left(\frac{r}{R_i}\right)^{n-1} n \bar{C}_{n,m}^i(r) \bar{Y}_{n,m}(\Omega)$$

r, Ω - spherical geocentric coordinates,

GM - planetocentric gravitational constant,

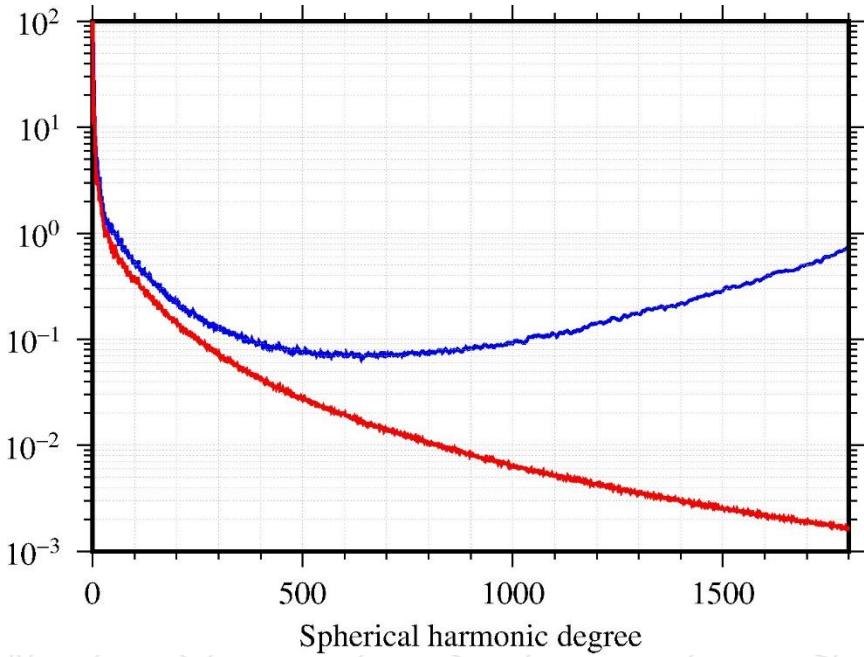
$\bar{Y}_{n,m}$ - spherical harmonic function of degree n and order m ,

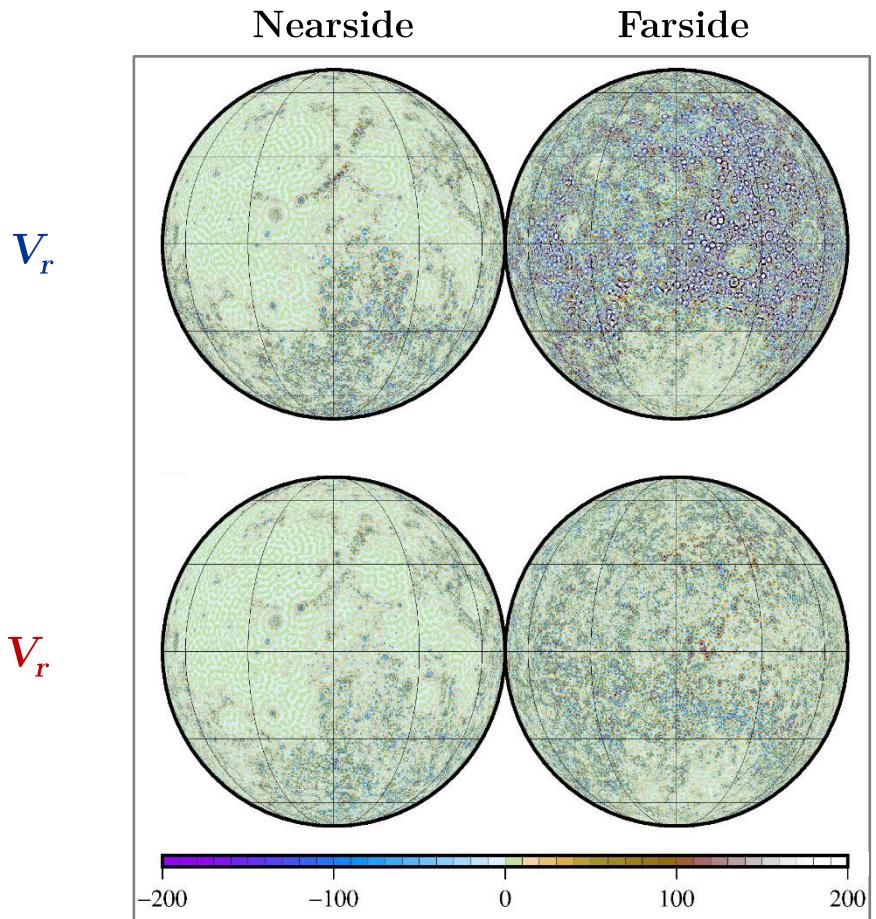
$\bar{C}_{n,m}$, $\bar{C}_{n,m}^o$, $\bar{C}_{n,m}^i$ - spherical harmonic coefficients of degree n and order m ,

R , R_o , R_i - scale factors,

N - maximum degree of the expansion.

Square-root of degree-order variances [mGal]:





Statistics (in mGal)

Model	Min.	Max.	Mean	Std. dev.
V_r	-1354.94	1333.11	-0.16	64.73
V_r	-248.83	274.95	-0.18	39.72

(Spherical harmonic spectrum 150-600)

Correlation coefficients

Quantity	Nearside	Farside	Global
V_r wrt. V_r	0.462	-0.135	0.050

4. Second-order radial derivative of V:

External SHE:

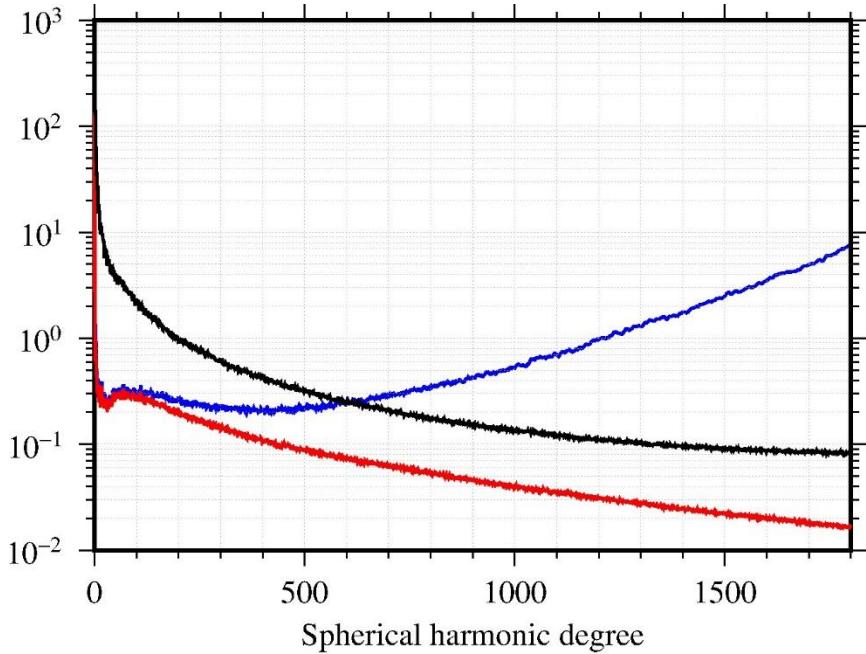
$$V_{rr}(r, \Omega) = \frac{GM}{R^3} \sum_{n,m}^N \left(\frac{R}{r}\right)^{n+3} (n+1)(n+2) \bar{C}_{n,m} \bar{Y}_{n,m}(\Omega)$$

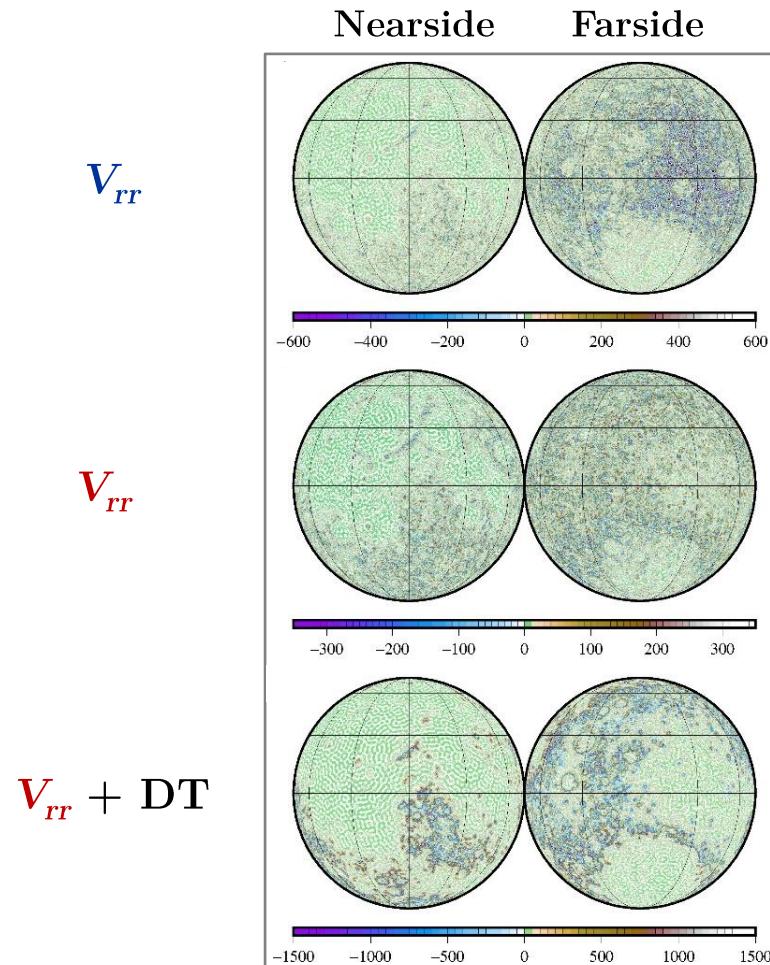
Internal SHE:

$$\begin{aligned} V_{rr}(r, \Omega) = & \frac{GM}{R_o^3} \sum_{n,m}^N \left(\frac{R_o}{r}\right)^{n+3} (n+1)(n+2) \bar{C}_{n,m}^o(r) \bar{Y}_{n,m}(\Omega) \\ & + \frac{GM}{R_i^3} \sum_{n,m}^N \left(\frac{r}{R_i}\right)^{n-2} n(n-1) \bar{C}_{n,m}^i(r) \bar{Y}_{n,m}(\Omega) \\ & - \frac{4\pi GM}{R_\varrho^3} \sum_{n,m}^N \bar{\varrho}_{n,m}(r) \bar{Y}_{n,m}(\Omega) \end{aligned}$$

$\bar{\varrho}_{n,m}$ - spherical harmonic coefficients of density,
 R_ϱ - scale factor.

Square-root of degree-order variances [E]:





Statistics (in E)

Quantity	Min.	Max.	Mean	Std. dev.
V_{rr}	-3753.14	3134.32	0.30	120.69
V_{rr}	-600.69	559.45	0.32	72.77
$V_{rr} + DT$	-2283.12	2458.96	0.09	306.39

(Spherical harmonic spectrum 150-600)

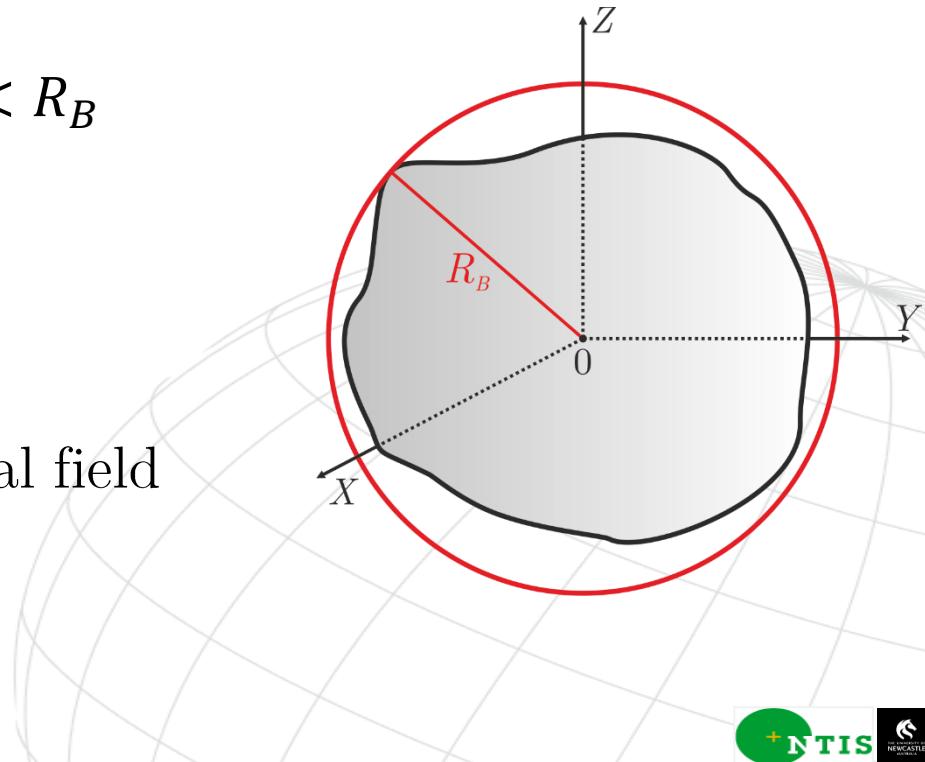
Correlation coefficients

Quantity	Nearside	Farside	Global
V_{rr} wrt. V_{rr}	0.927	0.737	0.778
$V_{rr} + DT$ wrt. V_{rr}	-0.348	-0.225	-0.266

5. Conclusions:

- Justification of the analytical continuation examined,
- Failure of the external SHEs @ $r < R_B$ demonstrated,
- Density cannot be neglected when calculating V_{rr} ,
- Implications for future gravitational field modelling and its applications.

<input checked="" type="checkbox"/>	$r > R_B$
<input type="checkbox"/>	$r < R_B$



Published article:

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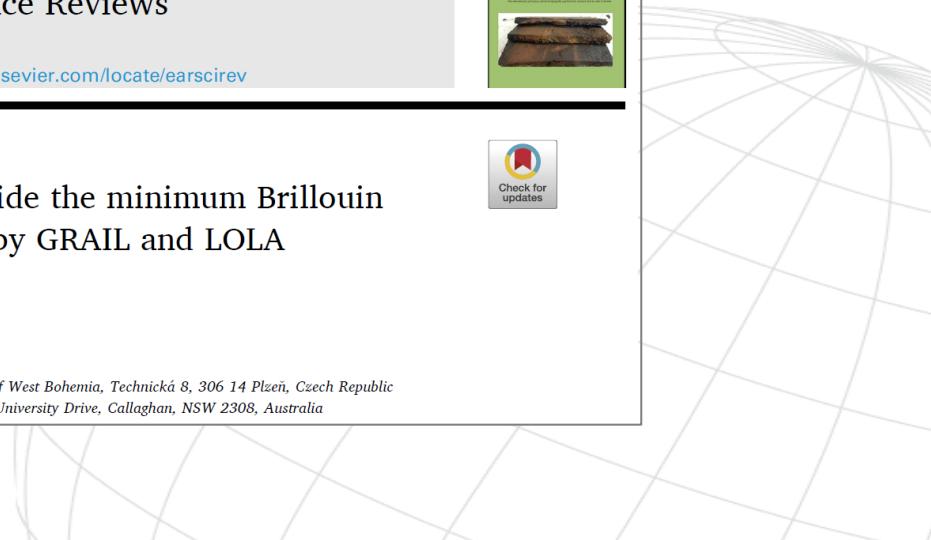
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On the use of spherical harmonic series inside the minimum Brillouin sphere: Theoretical review and evaluation by GRAIL and LOLA satellite data

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Thank you for your attention!

